The unitary-model-operator approach to nuclear many-body problems

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Abstract. Microscopic nuclear structure calculations have been performed within the framework of the unitary-model-operator approach. Ground-state and single-particle energies are calculated for nuclei around ¹⁴C, ¹⁶O and ⁴⁰Ca with modern nucleon-nucleon interactions.

1. Introduction

Recently, ab initio nuclear structure calculations starting with a realistic nucleon-nucleon (NN) interaction have been possible beyond few-nucleon systems. In this kind of study, the microscopic derivation of an effective interaction from the realistic NN interaction allows one to perform the structure calculation in a restricted model space. One of the successful methods of the structure calculation in this direction would be the no-core shell model (NCSM) [1, 2]. In the NCSM, the microscopic effective interaction (Hamiltonian) that takes account of the short-range correlation of the original NN interaction can be derived through a unitary transformation of the original Hamiltonian. Although the shell-model diagonalization is done in a large model space so that the final results do not depend on the model-space size, such a model space is still much smaller than the huge Hilbert space of the original Hamiltonian. The microscopic understanding of nuclear structure has been growing through the NCSM as well as the Green's function Monte Carlo (GFMC) [3, 4] in which the original NN interaction and a three-nucleon force can be directly used. However, although the GFMC and the NCSM are powerful method to describe nuclear structure, the application of these methods to the nuclear structure calculation may be limited to light nuclei up to $A \simeq 12$ due to the present computer power.

If one wishes to describe heavier nuclei, one needs to have another method. One of the promising methods may be the unitary-model-operator approach (UMOA) [5, 6, 7]. The UMOA can be regarded as one of the coupled-cluster methods (CCM) [8, 9, 10, 11, 12] of Hermitian type. As for recent developments of the CCM in nuclear theory, one may refer to Refs. [13, 14]. In the UMOA, an energy independent and Hermitian effective interaction is derived through a unitary transformation of the original Hamiltonian [15, 16] which is essentially the same as the unitary transformation used in the NCSM. By doing the unitary transformation in a two-step procedure, the structure calculation can be performed beyond p-shell nuclei. So far, we have

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performed structure calculations for not only stable nuclei around $^{16}{\rm O}$ [7] but also neutron-rich oxygen isotopes [17] and Λ hypernuclei [6]. In the following sections, we shall outline the calculation method of the UMOA and present recent results for nuclei around $^{14}{\rm C}$, $^{16}{\rm O}$ and $^{40}{\rm Ca}$.

2. Method of calculation

In the UMOA, the Hamiltonian \tilde{H} to be considered is obtained through a unitary transformation of the original Hamiltonian H as $\tilde{H}=e^{-S}He^{S}$. The exponent S is an anti-Hermitian operator and written as $S=\arctan(\omega-\omega^{\dagger})$ [5, 16] with a mapping operator $\omega=Q\omega P$ under the restrictive conditions PSP=QSQ=0, where P and Q are the usual projection operators and have the properties as P+Q=1, $P^2=P$, $Q^2=Q$ and PQ=QP=0. Thus, the operator ω satisfies the relation $\omega^2=\omega^{\dagger^2}=0$. We should note here that the unitary-transformation operator $U=e^S$ is also given by a block form concerning the P and Q spaces as

$$U = \begin{pmatrix} P(1+\omega^{\dagger}\omega)^{-1/2}P & -P\omega^{\dagger}(1+\omega\omega^{\dagger})^{-1/2}Q\\ Q\omega(1+\omega^{\dagger}\omega)^{-1/2}P & Q(1+\omega\omega^{\dagger})^{-1/2}Q \end{pmatrix}$$
(1)

which agrees with the unitary transformation by \bar{O} kubo [15]. By applying the above unitary transformation to a two-body subsystem of the original Hamiltonian, the two-body effective interaction \tilde{v}_{12} of Hermitian type is given by

$$\tilde{v}_{12} = U^{-1}(h_0 + v_{12})U - h_0, \tag{2}$$

where h_0 is the one-body part and v_{12} is the bare two-body interaction. The effective interaction in the P and, if necessary, Q spaces is determined by solving the decoupling equation $Q\tilde{v}_{12}P = P\tilde{v}_{12}Q = 0$. The actual method of calculating the matrix elements of U and \tilde{v}_{12} with the harmonic-oscillator (h.o.) basis states in the neutron-proton (np) formalism may be found in Ref. [7].

2.1. Two-step calculation for the effective interaction

In order to make the structure calculation in an inexpensive way, we perform the unitary transformation twice as follows. First, we derive the two-body effective interaction in a large model space to take into account the short-range correlation of the original NN interaction. The large model space consisting of two-body states is specified by a boundary number ρ_1 which is given with the sets of h.o. quantum numbers $\{n_a, l_a\}$ and $\{n_b, l_b\}$ of the two-body states by $\rho_1 = 2n_a + l_a + 2n_b + l_b$. The value of ρ_1 is taken as large as possible so that the calculated results do not depend on this value. If we diagonalize the transformed Hamiltonian with the many-body shell-model basis states using the effective interaction in the large model space, that leads to the NCSM. However, the calculations for heavier nuclei such as ¹⁶O and ⁴⁰O may not be practical in this manner because the present computer cannot diagonalize the huge matrix elements of the Hamiltonian which is large enough to obtain the converged results. If we intend to obtain only the energies of the ground states of closed-shell nuclei and the single-particle (-hole) states in its neighboring nuclei, it would be convenient to perform the unitary-transformation again as follows. We define a small model space $P_{np}^{(2)}$ and its $complement Q_{np}^{(2)}$ by separating the large model space in the previous procedure as shown in Fig. 1. The symbols ρ_n and ρ_p in Fig. 1 stand for the uppermost occupied states of the neutron and proton, respectively. We here only show the np channel. It should be noted that the $P_{np}^{(2)}$ and $Q_{np}^{(2)}$ spaces are considered on an equal footing in this second-step calculation when we solve the decoupling equation for the effective interaction $\tilde{v}_{12}^{(2)}$ in this step as $Q_{np}^{(2)}\tilde{v}_{12}^{(2)}P_{np}^{(2)}=0$ using the effective interaction determined in the first-step calculation as an input. Namely, we derive the effective interaction again in the large

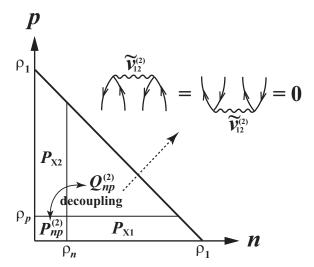


Figure 1. The spaces consisting of twobody states for the np channel considered in the second-step calculation for the effective interaction.

model space. However, by taking the model space $P_{np}^{(2)}$ and its complement $Q_{np}^{(2)}$ as shown in Fig. 1, the resultant effective interaction $\tilde{v}_{12}^{(2)}$ in this second step has no vertices which induce two-particle two-hole (2p2h) excitation. This is analogous to the Hartree-Fock (HF) condition which means that an original Hamiltonian is transformed so that the matrix elements for 1p1h excitation reduce to zero. Although the vertices of the one-body non-diagonal matrix elements remain in determining the effective interaction in the UMOA, these non-diagonal matrix elements are diagonalized at the end of the structure calculation. We notice that the two-body effective interaction in the first- and second-step calculations is determined self-consistently with the one-body potential for both the particle and hole states. Although this procedure is not necessarily needed in determining the effective interaction, such an effective interaction may be optimized for a restricted model space so as to obtain a good unperturbed energy.

We here show an example of the effect of renormalizing 2p2h excitation in the second-step calculation for the CD-Bonn potential [18]. The unperturbed ground-state energies of ¹⁶O for $\rho_1 = 12$ and $\hbar\Omega = 15$ MeV using the effective interactions determined in the first- and secondstep calculations are -38.40 and -104.12 MeV, respectively. This means that a large amount of the 2p2h effect is renormalized into the effective interaction in the small model space in the second-step calculation. In order to obtain the final result including the rest correlation effects, we diagonalize the transformed Hamiltonian with the shell-model basis states, taking into account 1p1h excitation from the unperturbed ground state. As for the closed-shell nucleus plus one-particle (one-hole) system, the shell-model basis states are composed of the 1p and 2p1hstates (1h and 1p2h states). These correlation energies are added to the unperturbed groundstate energies, and then we obtain the final results of the ground-state energies of the closed-shell nucleus and the closed-shell nucleus plus one-particle (one-hole) system. We note here that there remain some correction terms to be evaluated, such as the three-body cluster terms. Although the three-body cluster effect is essentially small, that may have a significant effect for some particular cases. Actually, we have found that the evaluation of the three-body cluster effect plays an important role to describe shell structure in neutron-rich oxygen isotopes [17]. However, generally speaking, the magnitude of the three-body cluster effect is much smaller than the twobody cluster effect, and thus the cluster expansion of the transformed Hamiltonian is justified from the viewpoint of the perturbative expansion. In the next section, we show some of the recent results which do not include the three-body cluster effect. The results including the three-body cluster effect will be reported elsewhere in the near future.

Table 1. The calculated ground-state energies with the 1p1h effect $E_{\rm g.s.}$ and the binding energies per nucleon BE/A of $^{16}{\rm O}$. All energies are in MeV.

¹⁶ O	Nijm 93	Nijm I	N^3LO	CD Bonn	Expt.
$E_{\rm g.s.}$ BE/A	-99.69 6.23	-104.25 6.52	-110.00 6.88	-115.61 7.23	-127.62 7.98

3. Results and discussion

First of all, we show calculated ground-state energies of 16 O with the 1p1h effect with modern NN interactions in Table 1. The final results for the Nijm-93, Nijm-I [19], the CD-Bonn [18] and the chiral N³LO [20] potentials are tabulated together with the experimental values. The binding energies per nucleon are also shown. The Coulomb force is used commonly. The results for the Nijm 93 and the CD Bonn are the least and most attractive, respectively, of the four potentials. This tendency can also be observed in the Faddeev-Yakubovsky calculations for 4 He by Nogga et al [21]. It is seen that the calculated ground-state energies are less bound than the experimental value. In the present results, higher-order correlations such as the three-body cluster terms have not been evaluated. If we include the effect of the three-body cluster terms, the calculated results become more attractive by a few MeV. For example, in the case of the

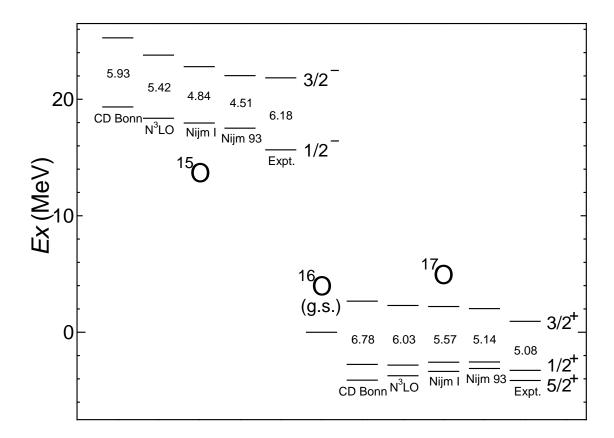
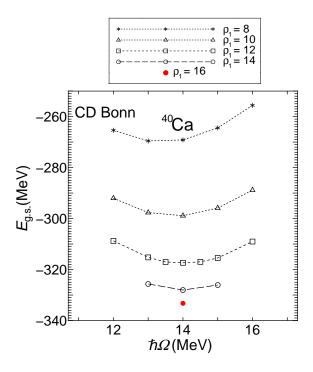


Figure 2. The calculated single-particle and single-hole energy levels in ¹⁷O and ¹⁵O, respectively, with modern NN interactions relative to the ground state of ¹⁶O.



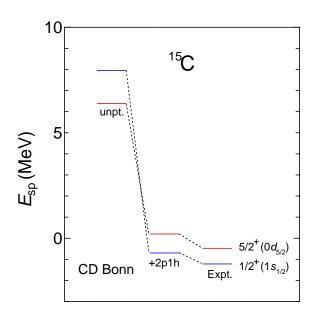


Figure 3. The ρ_1 and $\hbar\Omega$ dependences of the calculated ground-state energies of $^{40}\mathrm{Ca}$ for the CD-Bonn potential.

Figure 4. The calculated single-particle energies in ¹⁵C for the CD-Bonn potential.

CD-Bonn potential, the three-body cluster effect is about -4 MeV. However, even if we add this value to the ground-state energy in Table 1, the result is less attractive than the experimental value. In the present calculation, a genuine three-body force is not included. The inclusion of the real three-body force and the higher-order many-body correlations would compensate for the discrepancies between the experimental and calculated values. Such a study remains as an important task for a deeper understanding of nuclear ground-state properties in the UMOA.

In Fig. 2, calculated single-particle levels in ¹⁷O and single-hole levels in ¹⁵O relative to the ground state of ¹⁶O are illustrated. The values in Fig. 2 are the spin-orbit splitting energies for hole states in ¹⁵O and particle states in ¹⁷O. We see that the calculated spin-orbit splittings for the hole states in ¹⁵O are smaller than the experimental value. As for the spin-orbit splittings for the particle states ¹⁷O, as opposed to the hole-state case, the energies are larger than the experimental value. This may be due to an insufficient treatment for the 3/2⁺ unbound state in the present calculation. We have used only the h.o. states as the basis states. However, we may say that the calculated spin-orbit splittings for the hole and particle states in nuclei around ¹⁶O are, on the whole, not very different from the experimental values though the results somewhat depend on the interactions employed. The magnitudes of remaining discrepancies may be reduced if we include the genuine three-body force in the calculation and evaluate the higher-order cluster terms.

We are now trying to do calculations for heavier systems. Here we show some of the preliminary results for 40 Ca. Figure 3 shows the ρ_1 and $\hbar\Omega$ dependences of the ground-state energy with the 1p1h effect for the CD-Bonn potential. It is seen that a fairly convergent result is obtained at $\rho_1 = 16$ and $\hbar\Omega = 14$ MeV. Though we are now calculating for $\rho_1 = 18$, the difference of the results between $\rho_1 = 16$ and $\rho_1 = 18$ would be a few MeV at most. In the case of 16 O, the convergent result can be obtained at $\rho_1 = 14$. Since we are now calculating heavier

systems, we need a larger model space to obtain the convergent result.

In Fig. 4, we show calculated single-particle energies in a neutron-rich nucleus 15 C together with the experimental values. The result for "unpt." denotes the unperturbed single-particle energy which is defined as the h.o. kinetic energy plus the self-consistent one-body potential determined in the second-step calculation. In 15 C, the ordering of the experimental single-particle $1/2^+$ and $5/2^+$ states are opposite to the case of 17 O. However, our result for "unpt." does not reproduce this tendency, and the single-particle levels are rather repulsive to the experimental values. However, if we see the results with the 2p1h effect by the diagonalization as shown in the middle of Fig. 4, the results become more attractive and the two levels are reversed, and then a good agreement with the experimental value is obtained. We should remark, however, that the three-body cluster terms remain to be evaluated. As has been reported for neutron-rich oxygen isotopes in Ref. [17], the three-body cluster has a non-negligible effect on particle states of loosely-bound neutron-rich systems. Therefore, for the complete study, we have to evaluate the three-body cluster terms and also investigate the effect of the genuine three-body force though the present two-body cluster approximation should be a good treatment.

As has been shown before, the UMOA is a useful many-body theory to microscopically describe nuclear structure near closed-shell nuclei. The ground-state energy and the single-particle (-hole) energy can be calculated systematically for not only $N \simeq Z$ nuclei but also neutron-proton asymmetric systems beyond p-shell nuclei. Forthcoming new facilities for accelerating RI beams in the world will reveal new and exciting phenomena such as the change of shell structure in asymmetric nuclei. The UMOA has a possibility to develop new structures from a microscopic point of view in investigating the exotic nuclear systems.

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